# CONSTRAINTS FOR THE EXISTENCE OF FLAT METASTABLE NON-SUSY VACUA IN SUGRA

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- Viable SUSY breaking in SUGRA theories.
- Constraints for minimal chiral SUGRA models.
- Constraints for gauge invariant SUGRA models.
- Interplay between *F* and *D* breaking effects.
- Factorizable scalar manifolds.
- Symmetric scalar manifolds.
- Implications for string moduli.

# SUSY BREAKING AND SUGRA

In a viable SUGRA model, the vacuum state must be associated with a stationary point of the scalar potential where SUSY is spontaneously broken.

To get a realistic situation, there are however two additional conditions that must certainly be imposed:

- Flatness: The energy of the vacuum should be negligibly small, and reproduce the tiny value of the cosmological constant.
- Stability: The squared masses for small fluctuations around the vacuum should be positive.

The natural question is then whether these two conditions can be used to restrict the class of models of potential interest.

### MINIMAL SUGRA MODELS

A model with chiral multiplets  $\Phi^i$  is specified by a real Kähler potential Kand a holomorphic superpotential W. It has a Kähler symmetry for which  $(K, W) \rightarrow (K + F + \overline{F}, e^{-F}W)$ , and depends only on:

$$G=K+\log W+\log ar W$$

In the superconformal formulation, with a chiral compensator multiplet  $\Phi$ , the Kähler symmetry becomes manifest, with  $\Phi \to e^{F/3}\Phi$ . One can then set  $(K, W) \to (G, 1)$ , and write the Lagrangian in the form:

$$\mathcal{L} = \int\!\! d^4 heta \Big[ -3\,e^{-G/3} \Big] \, \Phi^\dagger \Phi + igg( \int\!\! d^2 heta\, \Phi^3 + ext{h.c.} igg)$$

The component action is obtained by freezing  $\Phi$  to gauge fix the extra conformal symmetries.

The scalar fields  $\phi^i$  behave as coordinates of a Kähler manifold, whose metric can be used to raise and lower chiral indices and is given by the second derivatives of *G*:

$$g_{i\bar{j}} = G_{i\bar{j}}$$

The auxiliary fields  $F^i$  are instead completely determined by the first derivatives of G:

$$F^i = -e^{G/2} \, G^i$$

The kinetic term is given by

$$T = g_{iar{j}}\,\partial_\mu \phi^i \partial^\mu \phi^{ar{j}}$$

The potential has instead the form:

$$V=e^G\Bigl(G^kG_k-3\Bigr)$$

Cremmer, Julia, Scherk, Ferrara, Girardello, Van Nieuwenhuizen Bagger, Witten

### Vacuum

The flatness condition is V = 0 and it implies that:

$$G^k G_k - 3 = 0$$

The stationarity conditions can be written as  $\nabla_i V = 0$  and imply:

$$G_i + G^k \nabla_i G_k = 0$$

The stability conditions amount finally to imposing

$$egin{pmatrix} m_{ij}^2 & m_{ij}^2 \ m_{ij}^2 & m_{ij}^2 \end{pmatrix} > 0$$

where the blocks  $m_{i\bar{j}}^2 = \nabla_i \nabla_{\bar{j}} V$  and  $m_{ij}^2 = \nabla_i \nabla_j V$  are given by:

$$m_{i\overline{j}}^2 = e^G \Big[ g_{i\overline{j}} + 
abla_i G_k 
abla_{\overline{j}} G^k - R_{i\overline{j}p\overline{q}} G^p G^{\overline{q}} \Big]$$
  
 $m_{ij}^2 = e^G \Big[ 2 \, 
abla_{(i}G_{j)} + G^k 
abla_{(i} 
abla_{j)} G_k \Big]$ 

Supersymmetry is spontaneously broken, and the gravitino mass is:

$$m_{3/2} = e^{G/2}$$

The would-be Goldstino fermion is identified with the linear combination  $\eta = f_i \psi^i$ , where:

$$f_i = rac{1}{\sqrt{3}} rac{F_i}{m_{3/2}} = -rac{1}{\sqrt{3}} \, G_i$$

For fixed  $g_{ij}$ , the quantities  $f_i$  can be treated as independent variables without any particular constraint.

### Flatness condition

The flatness condition is the constraint that the Goldstino vector should have unit length:

$$g_{iar{\jmath}}\,f^if^{ar{\jmath}}=1$$

### Stability condition

The stability condition is more complicated and can be studied only model by model, by explicit diagonalization.

It is however possible to find simpler but weaker conditions for stability, which are necessary but not sufficient, by looking at particular directions in scalar field space.

In this case, there is only one special complex direction that appears in the problem and that we could use: the Goldstino direction  $G^i$ .

Looking at the two independent real directions  $(G^i, G^{\vec{i}})$  and  $(iG^i, -iG^{\vec{i}})$ , one deduces the condition  $m_{i\bar{j}}^2 G^i G^{\vec{j}} > 0$ . Using the flatness and the stationarity conditions, this gives:

$$R_{iar{\jmath}par{q}}\,f^if^{ar{\jmath}}f^pf^{ar{q}} < rac{2}{3}$$

### Constraints

Summarizing, a stationary point can lead to a satisfactory situation only if two simple flatness and stability conditions are satisfied at it.

It is convenient redefine the fields to locally switch to flat indices, with  $g_{I\bar{J}} = \delta_{I\bar{J}}$ . The two conditions become then simply:

Flatness: 
$$\delta_{I\bar{J}} f^I f^{\bar{J}} = 1$$
  
Stability:  $R_{I\bar{J}P\bar{Q}} f^I f^{\bar{J}} f^P f^{\bar{Q}} < \frac{2}{3}$ 

The flatness condition fixes the overall amount of SUSY breaking. The stability condition requires the existence of directions with R < 2/3, and constrains the direction of SUSY breaking to be close to these.

Gomez-Reino, Scrucca

### GAUGE INVARIANT SUGRA MODELS

A model with chiral multiplets  $\Phi^i$  and vector multiplets  $V^a$  is specified by a real Kähler function G and a holomorphic gauge kinetic matrix  $H_{ab}$ .

In the superconformal formulation, the Lagrangian has the form:

$$egin{split} \mathcal{L} &= \int\!\!d^4 heta \Big[\! -3\,e^{-G/3}\Big]\,\Phi^\dagger\Phi + \left(\,\int\!\!d^2 heta\,\Phi^3 + \mathrm{h.c.}
ight) \ &+ \left(\,\int\!\!d^2 heta\,rac{1}{4}H_{ab}W^{alpha}W^{blpha}_{lpha} + \mathrm{h.c.}
ight) \end{split}$$

The gauge transformations of the superfields are encoded in a set of holomorphic Killing vectors  $X_a^i$ :

$$\delta\Phi^i=\Lambda^a X^i_a \qquad \delta V^a=-i(\Lambda^a-ar\Lambda^a)$$

The local charges are:

$$Q_{ai}{}^j = -\nabla_i X_a^j$$

The function G must be invariant:  $\delta G = 0$ . This implies:

$$G_a = -i X_a^i G_i = i X_a^{\overline{i}} G_{\overline{i}}$$

The first and second derivatives of these relations imply that:

The function  $H_{ab}$  must instead transform in such a way to cancel possible residual quantum anomalies:  $\delta H_{bc} = i \Lambda^a A_{abc}$ . This implies:

$$X_a^i \nabla_i H_{bc} = i \, A_{abc}$$

The scalar fields  $\phi^i$  parametrize now a symmetric Kähler manifold. The metric for chiral indices is given by the second derivatives of G:

$$g_{i\overline{j}} = G_{i\overline{j}}$$

The vector fields  $A^a_{\mu}$  gauge the symmetries associated to the isometries  $X^i_a$ . The real part of  $H_{ab}$  effectively acts as a metric for vector indices, while its imaginary part gives additional parameters:

$$h_{ab} = \operatorname{Re} H_{ab} \qquad \quad \theta_{ab} = \operatorname{Im} H_{ab}$$

The auxiliary fields  $F^i$  and  $D^a$  are given by the first derivatives of G:

$$F^i = -e^{G/2} G^i \qquad D^a = -G^a$$

The kinetic terms are:

$$T=g_{iar{\jmath}}\left(\partial_{\mu}\phi^{i}-X^{i}_{a}A^{a}_{\mu}
ight)\left(\partial^{\mu}\phi^{ar{\imath}}-X^{ar{\imath}}_{a}A^{a\mu}
ight) 
onumber \ -rac{1}{4}h_{ab}\,F^{a}_{\mu
u}F^{b\mu
u}-rac{1}{4} heta_{ab}\,F^{a}_{\mu
u} ilde{F}^{b\mu
u}$$

The potential has instead the form:

$$V=e^G \Big(G^k G_k-3\Big)+rac{1}{2}\,G^a G_a$$
 .

Cremmer, Ferrara, Girardello, Van Proeyen Bagger

### Vacuum

The flatness condition is V = 0 and it implies that:

$$G^{k}G_{k} + \frac{1}{2}e^{-G}G^{a}G_{a} - 3 = 0$$

The stationarity conditions can be written as  $\nabla_i V = 0$  and imply:

$$G_i + G^k 
abla_i G_k + e^{-G} \Big[ G^a \Big( 
abla_i - rac{1}{2} G_i \Big) G_a + rac{1}{2} h_{abi} G^a G^b \Big] = 0$$

The stability conditions amount in this case to imposing the slightly weaker requirement:

$$egin{pmatrix} m_{ij}^2 & m_{ij}^2 \ m_{\overline{ij}}^2 & m_{\overline{ij}}^2 \end{pmatrix} \geq 0$$

The equality sign holds for the would-be Goldstone bosons, which are absorbed by the vector fields.

The blocks  $m_{i\bar{j}}^2 = \nabla_i \nabla_{\bar{j}} V$  and  $m_{ij}^2 = \nabla_i \nabla_j V$  are now given by the following more complicated expressions:

$$\begin{split} m_{i\overline{j}}^2 &= e^G \Big[ g_{i\overline{j}} - R_{i\overline{j}p\overline{q}} G^p G^{\overline{q}} + \nabla_i G_k \nabla_{\overline{j}} G^k \Big] \\ &+ \Big[ \frac{1}{2} \Big( G_i G_{\overline{j}} - g_{i\overline{j}} \Big) G^a G_a + \Big( G_{(i}h_{ab\overline{j})} + h^{cd}h_{aci}h_{bd\overline{j}} \Big) G^a G^b \\ &- 2 \, G^a G_{(i} \nabla_{\overline{j})} G_a - 2 \, G^a h^{bc}h_{ab(i} \nabla_{\overline{j})} G_c \\ &+ h^{ab} \nabla_i G_a \nabla_{\overline{j}} G_b + G^a \nabla_i \nabla_{\overline{j}} G_a \Big] \\ m_{ij}^2 &= e^G \Big[ 2 \, \nabla_{(i}G_{j)} + G^k \nabla_{(i} \nabla_{j)} G_k \Big] \\ &+ \Big[ \frac{1}{2} \Big( G_i G_j - \nabla_{(i}G_{j)} \Big) G^a G_a + \Big( G_{(i}h_{abj)} + h^{cd}h_{aci}h_{bdj} \Big) G^a G^b \\ &- \frac{1}{2} h_{abij} G^a G^b - 2 G^a G_{(i} \nabla_{j)} G_a - 2 \, G^a h^{bc}h_{ab(i} \nabla_{j)} G_c \\ &+ h^{ab} \nabla_i G_a \nabla_j G_b + G^a \nabla_i \nabla_{\overline{j}} G_a \Big] \end{split}$$

Supersymmetry is spontaneously broken, and the gravitino mass has the same expression as before:

$$m_{3/2} = e^{G/2}$$

The would-be Goldstino fermion is  $\eta = f_i \psi^i + d_a \lambda^a$ , where:

$$f_i = \frac{1}{\sqrt{3}} \frac{F_i}{m_{3/2}} = -\frac{1}{\sqrt{3}} G_i \qquad d_a = \frac{1}{\sqrt{6}} \frac{D_a}{m_{3/2}} = -\frac{1}{\sqrt{6}} e^{-G} G_a$$

Gauge symmetries are also spontaneously broken, and the vector mass matrix is:

$$M^2_{ab} = 2 \, g_{iar{j}} X^i_a X^{ar{j}}_b = 2 \, g_{iar{j}} \, 
abla^i G_a 
abla^{ar{j}} G_b$$

The would-be Goldstone bosons are  $\sigma_a = v_{ai}\phi^i + v_{a\bar{\imath}}\phi^{\bar{\imath}}$ , where:

$$v_{ai} = rac{X_{ai}}{\sqrt{X_a^k X_{ak}}}$$

For fixed  $g_{i\bar{j}}$ ,  $X_a^i$  and  $h_{ab}$ , the quantities  $f_i$  and  $d_a$  can be thought as variables, but with some relations involving the parameters:

$$x_a^i = rac{X_a^i}{m_{3/2}} \qquad m_{ab} = rac{1}{2} rac{M_{ab}}{m_{3/2}} \qquad q_{aiar{\jmath}} = rac{Q_{aiar{\jmath}}}{m_{3/2}} \qquad a_{abc} = rac{A_{abc}}{m_{3/2}}$$

There is a dynamical relation holding at stationary points, which is implied by stationarity along the directions  $X_a^i$ :

$$q_{aij} f^i f^j - \sqrt{rac{2}{3}} \left[ 2 \, m_{ab}^2 + \left( 3 f^i f_i - 1 
ight) h_{ab} 
ight] d^b + a_{abc} \, d^b d^c = 0$$
 Kawamura

There is then a kinematical relation holding at any point, which is implied by gauge invariance of G:

$$d_a = -\frac{i}{\sqrt{2}} x_a^i f_i = \frac{i}{\sqrt{2}} x_a^{\vec{\imath}} f_{\vec{\imath}}$$

There is finally a kinematical bound, implied by this relation:

$$|d_a| \leq m_{aa} \sqrt{f^i f_i}$$

### Flatness condition

The flatness condition is again simply the constraint that the Goldstino vector should have unit length:

$$g_{i\bar\jmath}\,f^if^{\bar\jmath}+h_{ab}\,d^ad^b=1$$

Stability condition

The stability condition is as before a complicated condition, which can be studied only model by model, by explicit diagonalization.

However, once again it is possible to find simpler but weaker conditions for stability, which are necessary but not sufficient, by looking at particular directions in scalar field space.

In this case, there are two kinds of special complex directions that appear: the projected Goldstino direction  $G^i$  and the Goldstone directions  $X_a^i$ . Looking at the real directions  $(G^i, G^{\vec{i}})$  and  $(iG^i, -iG^{\vec{i}})$ , one deduces the condition  $m_{i\vec{j}}^2 G^i G^{\vec{j}} \ge 0$ . Using the flatness and the stationarity conditions, this yields:

$$\begin{split} R_{i\bar{j}p\bar{q}} \, f^i f^{\bar{j}} f^p f^{\bar{q}} + 2 \left( h_{ab} h_{cd} - \frac{1}{2} \, h_{ab}^{\phantom{ab}i} h_{cdi} \right) d^a d^b d^c d^d \\ &- 2 \, h^{cd} h_{aci} h_{bd\bar{j}} \, f^i f^{\bar{j}} \, d^a d^b + \sqrt{\frac{3}{2}} \, a_{abc} \, d^a d^b d^c \\ &- \frac{8}{3} \left( m_{ab}^2 - \frac{1}{2} \, h_{ab} \right) d^a d^b \leq \frac{2}{3} \end{split}$$

Looking at the real directions  $(X_a^i, X_a^{\vec{\imath}})$  and  $(iX_a^i, -iX_a^{\vec{\imath}})$ , one finds that the former are flat directions whereas the latter imply the extra conditions  $m_{ij}^2 X_a^i X_a^j \ge 0$ , which have however a complicated form.

No extra useful condition

### Constraints

As before, a stationary point can lead to a satisfactory situation only if two simple flatness and stability conditions are satisfied at it.

It is convenient redefine the fields to locally switch to flat indices, with  $g_{I\bar{J}} = \delta_{I\bar{J}}$  and  $h_{AB} = \delta_{AB}$ . For simplicity, we also assume a constant and diagonal gauge kinetic function. The two conditions read then:

$$\begin{array}{l} \mbox{Flatness: } \delta_{I\bar{J}}\,f^{I}f^{\bar{J}} = 1 - \sum_{A}d_{A}^{2} \\ \mbox{Stability: } R_{I\bar{J}P\bar{Q}}\,f^{I}f^{\bar{J}}f^{P}f^{\bar{Q}} \leq \frac{2}{3} + \frac{8}{3} \sum_{A} \Bigl(m_{A}^{2} - \frac{1}{2}\Bigr)d_{A}^{2} - 2\Bigl(\sum_{A}d_{A}^{2}\Bigr)^{2} \end{array}$$

The flatness condition fixes as before the amount of SUSY breaking. The stability condition constrains instead the directions of SUSY breaking.

The  $f_I$  represent the basic qualitative seed for SUSY breaking, whereas the  $d_A$  provide additional quantitative effects.

Indeed, the  $d_A$  are not independent from the  $f_I$ , but rather related to them as follows:

Dynamical relation: 
$$d_A = \sqrt{\frac{3}{8}} \frac{q_{AI\bar{J}} f^I f^J}{m_A^2 - \frac{1}{2} + \frac{3}{2} f^I f_I}$$
  
Kinematical relation:  $d_A = -\frac{i}{\sqrt{2}} x_A^I f_I$   
Kinematical bound:  $|d_A| \le m_A \sqrt{f^I f_I}$ 

These **3** relations are gradually weaker and simpler, and can be used to set up **3** different types of analyses of the constraints.

The effect of vector multiplets is generically to alleviate the constraints and results in a lowering of the effective curvature for chiral multiplets. One needs then  $\tilde{R} < 3/2$ , which is a milder constraint.

Gomez-Reino, Scrucca

### RELATIVE EFFECT OF F AND D BREAKING

It is useful to introduce the new variables:

$$ilde{f}^I = rac{f^I}{\sqrt{1-\sum_B d_B^2}} \qquad ilde{d}_A = rac{d_A}{\sqrt{1-\sum_B d_B^2}}$$

The constraints can then be rewritten as:

Flatness: 
$$\delta_{I\bar{J}} \tilde{f}^{I} \tilde{f}^{\bar{J}} = 1$$
  
Stability:  $R_{I\bar{J}P\bar{Q}} \tilde{f}^{I} \tilde{f}^{\bar{J}} \tilde{f}^{P} \tilde{f}^{\bar{Q}} \leq \frac{2}{3} \left(1 + \Delta(\tilde{d}_{A})\right)$ 

where:

$$\Delta( ilde{d}_A) = 4\sum_A m_A^2\, ilde{d}_A^2 + 4\sum_A \Bigl(m_A^2-1\Bigr) ilde{d}_A^2\sum_B ilde{d}_B^2$$

The dynamical relation between auxiliary fields becomes:

$$ilde{d}_A \sqrt{1\!+\!\sum_B \! ilde{d}_B^2} \left[1\!+\!m_A^2\!-\!rac{3}{2} rac{\sum_B \! ilde{d}_B^2}{1\!+\!\sum_B \! ilde{d}_B^2}
ight] = \sqrt{rac{3}{8}} \, q_{AIar{J}} \, ilde{f}^I ilde{f}^{ar{J}}$$

The kinematical relation becomes instead:

$$ilde{d}_A = -rac{i}{\sqrt{2}} \, x^I_A \, ilde{f}_I$$

Finally the kinematical bound implies:

$$| ilde{d}_A| \leq m_A$$

Exploiting these relations to eliminate the  $\tilde{d}_A$ , one can arrive to constraints involving only the  $\tilde{f}_I$  and a corrected effective curvature  $\tilde{R}_{I\bar{J}P\bar{Q}}$ , taking the simple form:

Flatness: 
$$\delta_{I\bar{J}} \tilde{f}^{I} \tilde{f}^{\bar{J}} = 1$$
  
Stability:  $\tilde{R}_{I\bar{J}P\bar{Q}} \tilde{f}^{I} \tilde{f}^{\bar{J}} \tilde{f}^{P} \tilde{f}^{\bar{Q}} \leq \frac{2}{3}$ 

### Exploiting the dynamical relation

Whenever  $|\tilde{d}_A| \ll 1$ , as for instance when  $m_A \gg 1$  or  $m_A \ll 1$ , the dynamical relation can be linearized and:

$$ilde{d}_A\simeq \sqrt{rac{3}{8}}\,rac{1}{1\!+\!m_A^2}\,q_{AIar{J}}\, ilde{f}^I ilde{f}^{ar{J}}$$

Moreover, keeping only the leading term in  $\Delta$ , one finds:

$$\Delta( ilde{d}_A)\simeq rac{3}{2}\sum_A iggl[rac{m_A}{1\!+\!m_A^2}iggr]^2 iggl| q_{AIar{J}}\, ilde{f}^I ilde{f}^{ar{J}}iggr|^2$$

The net effect of vector multiplets is then to change the effective curvature for chiral multiplets to:

$$\tilde{R}_{I\bar{J}P\bar{Q}} \simeq R_{I\bar{J}P\bar{Q}} - \sum_{A} \left[ \frac{m_A}{1+m_A^2} \right]^2 q_{AI(\bar{J}} q_{AP\bar{Q})}$$

Whenever  $|\tilde{d}_A| \sim 1$ , as generically for  $m_A \sim 1$ , one can combine the dynamical relation and the kinematical bound to derive the upper bound:

$$| ilde{d}_A|\sqrt{1\!+\!\sum_B\! ilde{d}_B^2} \le \sqrt{rac{3}{8}}\,rac{1\!+\!\sum_B m_B^2}{1\!+\!m_A^2\!+\!\left(m_A^2\!-\!rac{1}{2}
ight)\!\sum_B\!m_B^2}\,\Big|q_{AIar{J}}\, ilde{f}^I ilde{f}^{ar{J}}\Big|$$

Dropping the negative term in  $\Delta$ , one finds then:

$$\Delta( ilde{d}_A) \leq rac{3}{2} \sum_A igg[ rac{m_A \Big( 1 + \sum_B m_B^2 \Big)}{1 + m_A^2 + \Big( m_A^2 - rac{1}{2} \Big) \sum_B m_B^2} \Big]^2 \Big| q_{AIar{J}} \, ilde{f}^I ilde{f}^{ar{J}} \Big|^2$$

This can be used to get a simpler but weaker form of the constraints, where the net effect of vector multiplets is encoded in:

$$ilde{R}_{Iar{J}Par{Q}} \simeq R_{Iar{J}Par{Q}} - \sum_{A} \Biggl[ rac{m_A \Bigl(1 + \sum_B m_B^2 \Bigr)}{1 + m_A^2 + \Bigl(m_A^2 - rac{1}{2}\Bigr) \sum_B m_B^2} \Biggr]^2 q_{AI(ar{J}\, q_{APar{Q}})}$$

### Exploiting the kinematical relation

Useful results only when the theory has special properties.

### Exploiting the kinematical bound

Using the kinematical bound, one can derive an upper bound for  $\Delta$ . This is a complicated function of the  $m_A^2$ . But it can also be further bounded by above by a simple function of  $m^2 = \sum_A m_A^2$ :

$$\Delta\big(\tilde{d}_A\big) \leq \frac{m^4}{1\!-\!m^2}\,\theta\Big(\frac{1}{2}\!-\!m^2\Big) + 4\,m^6\,\theta\Big(m^2\!-\!\frac{1}{2}\Big)$$

This can be used to get a simpler but weaker form of the constraints, where the net effect of vector multiplets is encoded in:

$$ilde{R}_{Iar{J}Par{Q}} \simeq \left[1\!+\!rac{m^4}{1\!-\!m^2}\, heta\!\left(\!rac{1}{2}\!-\!m^2
ight)\!+\!4\,m^6\, heta\!\left(\!m^2\!-\!rac{1}{2}
ight)\!
ight]^{-1}\!R_{Iar{J}Par{Q}}$$

## FACTORIZABLE SPACES

Suppose that  $\mathcal{M}$  is a product of N distinct 1-dimensional manifolds. The function K splits then into a sum of terms depending on a single field, while W can instead still be arbitrary:

In this situation,  $g_{i\bar{j}}$  and  $R_{i\bar{j}p\bar{q}}$  are diagonal and have only N non-zero components. Moreover, these are related by  $R_{i\bar{\imath}i\bar{\imath}} = R_i g_{i\bar{\imath}}^2$ , in terms of the curvature scalars  $R_i$  of the various submanifolds.

The flat-index curvature tensor entering the constraints takes then the following form:

$$R_{Iar{J}Par{Q}} = \left\{egin{array}{cc} R_i \ , & ext{if } I = J = P = Q \ 0 \ , & ext{otherwise} \end{array}
ight.$$

This simplifies enough to problem to allow for an exact solution.

The flatness and stability conditions simplify to

Flatness: 
$$\sum_{k} \Theta_{k}^{2} = 1$$
  
Stability:  $\sum_{k} R_{k} \Theta_{k}^{4} < \frac{2}{3}$ 

in terms of the N real and positive variables

$$\Theta_i = |f_I|$$

These constraints admit solutions only if the curvatures satisfy

$$\sum_k R_k^{ extsf{-1}} > rac{3}{2}$$

The allowed breaking directions fill then a cone around the axis

$$\Theta^0_i = \sqrt{rac{R_i^{-1}}{\sum_k R_k^{-1}}}$$

More precisely, one finds  $\Theta_i \in [\Theta_i^-, \Theta_i^+]$ , with:

$$\begin{split} \Theta_{i}^{+} &= \begin{cases} \sqrt{\frac{R_{i}^{\text{-1}} + \sqrt{\frac{2}{3}R_{i}^{\text{-1}} \left(\sum_{k \neq i} R_{k}^{\text{-1}}\right) \left(\sum_{k} R_{k}^{\text{-1}} - \frac{3}{2}\right)}}{\sum_{k} R_{k}^{\text{-1}}}, & R_{i}^{\text{-1}} < \frac{3}{2} \\ 1 & , & R_{i}^{\text{-1}} > \frac{3}{2} \end{cases} \\ \Theta_{i}^{-} &= \begin{cases} \sqrt{\frac{R_{i}^{\text{-1}} - \sqrt{\frac{2}{3}R_{i}^{\text{-1}} \left(\sum_{k \neq i} R_{k}^{\text{-1}}\right) \left(\sum_{k} R_{k}^{\text{-1}} - \frac{3}{2}\right)}}{\sum_{k} R_{k}^{\text{-1}}}, & \sum_{k \neq i} R_{k}^{\text{-1}} < \frac{3}{2} \\ 0 & , & \sum_{k \neq i} R_{k}^{\text{-1}} > \frac{3}{2} \end{cases} \end{cases} \end{split}$$

The relevance of each  $\Phi_i$  for SUSY breaking depends thus on the size of the corresponding  $R_i^{-1}$  with respect to 3/2.

## SYMMETRIC SPACES

Suppose that  $\mathcal{M}$  is a coset space G/H, where G is a group of global isometries and H a local stability group. The function K has then some special form, but W can still be arbitrary:

In this situation,  $g_{i\bar{j}}$  and  $R_{i\bar{j}p\bar{q}}$  are *G*-invariant and there are relations among their components. Moreover  $R_{i\bar{j}p\bar{q}}$  can always be related to  $g_{i\bar{j}}$ , through some overall curvature scale  $R_{\rm all}$ .

Calabi, Vesentini

The flat-index curvature tensor entering the constraints has the special structure

$$R_{I\bar{J}P\bar{Q}} = R_{\mathrm{all}} \Big( G$$
-invariant combination of  $H$ -invariant  $\delta$ 's  $\Big)$ 

This again simplifies enough to problem to allow for an exact solution.

#### Generalized spheres

With N = 1 + q fields  $\Phi_i$  one can have:

$$\mathcal{M} = rac{SU(1,1+q)}{U(1) imes SU(1+q)} \supset rac{SU(1,1)}{U(1)}$$

The Riemann tensor in normal coordinates reads

$$R_{Iar{J}Par{Q}} = rac{1}{2} R_{ ext{all}} \Big( \delta_{Iar{J}} \, \delta_{Par{Q}} + \delta_{Iar{Q}} \, \delta_{Par{J}} \Big)$$

The flatness and stability conditions can then be rewritten as

Flatness: 
$$\Theta^2 = 1$$
  
Stability:  $R_{\rm all} \, \Theta^4 < {2 \over 3}$ 

in terms of just 1 real and positive variable

$$\Theta = \sqrt{\sum_k |f_K|^2}$$

The situation is then as for 1 field with  $R = R_{\rm all} \Rightarrow R_{\rm all}^{-1} > 3/2$ .

### Unitary Grassmannians

With N = p (p + q) fields  $\Phi_{ia}$  one can have:

$$\mathcal{M} = rac{SU(p,p+q)}{U(1) imes SU(p) imes SU(p+q)} \supset \left(rac{SU(1,1)}{U(1)}
ight)^p$$

The Riemann tensor in normal coordinates reads

$$R_{IA\,ar{J}ar{B}\,PC\,ar{Q}ar{D}} = rac{1}{2} R_{
m all} \Big( \delta_{Iar{J}}\,\delta_{Par{Q}}\,\delta_{Aar{D}}\,\delta_{Car{B}} + \delta_{Iar{Q}}\,\delta_{Par{J}}\,\delta_{Aar{B}}\,\delta_{Car{D}} \Big)$$

The flatness and stability conditions reduce then simply to

Flatness: 
$$\sum_{k} \Theta_{k}^{2} = 1$$
  
Stability:  $\sum_{k} R_{all} \Theta_{k}^{4} < \frac{2}{3}$ 

in terms of the p real and positive variables

$$\Theta_{i} = \left| \mathrm{Eigenvalue}_{i}(f_{IA}) \right|$$

The situation is then as for p fields with  $R_i = R_{all} \Rightarrow R_{all}^{-1} > 3/(2p)$ .

### **Orthogonal Grassmannians**

With N = 2 + q fields  $\Phi_i$  one can have:

$$\mathcal{M} = rac{SO(2,2+q)}{SO(2) imes SO(2+q)} \supset \left(rac{SU(1,1)}{U(1)}
ight)^2$$

The Riemann tensor in normal coordinates reads

$$R_{Iar{J}Par{Q}} = rac{1}{2} R_{ ext{all}} \Big( \delta_{Iar{J}} \, \delta_{Par{Q}} + \delta_{Iar{Q}} \, \delta_{Par{J}} - \delta_{IP} \, \delta_{ar{J}ar{Q}} \Big)$$

The flatness and stability conditions reduce then to

Flatness: 
$$\Theta_+^2 + \Theta_-^2 = 1$$
  
Stability:  $R_{\rm all} \left( \Theta_+^4 + \Theta_-^4 \right) < rac{2}{3}$ 

in terms of the 2 real and positive variables

$$\Theta_{\pm} = rac{1}{\sqrt{2}} \sqrt{\sum_k |f_K|^2 \pm \sqrt{\left(\sum_k |f_K|^2
ight)^2 - \left|\sum_k (f_K)^2
ight|^2}}$$

The situation is then as for 2 fields with  $R_i = R_{\rm all} \Rightarrow R_{\rm all}^{-1} > 3/4$ .

## MODULI IN STRING MODELS

In string models, a natural candidate for the breaking sector is that of the moduli  $M_i$  controlling the coupling strength and the compactification geometry, and the Wilson lines  $Z_a$  of the hidden gauge groups.

Kaplunovky, Louis

In the simplest models, the scalar manifold is factorizable and symmetric, because it emerges as a projection of a D = 10 SUSY theory. For a gauge group of rank s, this has the form

$$\mathcal{M} = rac{SU(1,1)}{U(1)} imes rac{SO(6,6+s)}{SO(6) imes SO(6+s)} igg|_{ ext{proj}}$$

Narain

The first factor describes the dilaton modulus S, and is always present. The second factor is spanned by the Kähler and complex structure moduli  $T_p$  and  $U_q$ , as well as the Wilson lines  $Z_a$ , and depends on the projection.

#### Minimal modulus geometry

The minimal factor that can arise for a single modulus  $M_i$  is described by the following Kähler potential:

$$K_i = -\,n_i \ln \Bigl(M_i + M_i^\dagger \Bigr)$$

Witten

This corresponds to the simplest possible symmetric space:

$$\mathcal{M}_i = rac{SU(1,1)}{U(1)}$$

This is the basic one-dimensional space with constant curvature that can appear, and the curvature scalar is given by:

$$R_i = rac{2}{n_i}$$

#### Maximally symmetric enhancement

Certain moduli  $M_i$  can mix to some number  $q_i$  of related Wilson lines  $Z_{a_i}$ , and these  $1 + q_i$  fields have then

$$K_i = -n_i \ln \left( M_i + M_i^\dagger - \sum_{a_i} Z_{a_i}^\dagger Z_{a_i} 
ight)$$

Ellis, Kounnas, Nanopoulos Ferrara, Kounnas, Porrati

The corresponding scalar manifold is given by:

$$\mathcal{M}_i = rac{SU(1,1+q_i)}{U(1) imes SU(1+q_i)} \supset rac{SU(1,1)}{U(1)}$$

This behaves as 1 copy of the minimal geometry for the flatness and stability constraints, with curvature scale:

$$R_i = rac{2}{n_i}$$

#### Unitary enhancement

A set of  $p_r$  moduli with the same  $n_r$  can get enhanced to  $p_r^2$  moduli  $M_{i_r j_r}$  and couple to some number  $p_r q_r$  of related Wilson lines  $Z_{i_r a_r}$ . These  $p_r(p_r + q_r)$  fields have then

$$K_r = -n_r \ln \det \left( M_{i_r j_r} + M^{\dagger}_{i_r j_r} - \sum_{a_r} Z^{\dagger}_{i_r a_r} Z_{j_r a_r} 
ight)$$

Ferrara, Kounnas, Porrati

The corresponding scalar manifold is:

$$\mathcal{M}_r = rac{SU(p_r,p_r+q_r)}{U(1) imes SU(p_r) imes SU(p_r+q_r)} \supset \left(rac{SU(1,1)}{U(1)}
ight)^{p_r}$$

This behaves as  $p_r$  copies of the minimal geometry for the flatness and stability constraints, with overall curvature scale given by:

$$R_r = \frac{2}{n_r}$$

#### Orthogonal enhancement

A pair of 2 moduli  $M_{1_r}$  and  $M_{2_r}$  with the same  $n_r$  can also mix in a more peculiar fashion to some number  $q_r$  of related Wilson lines  $Z_{a_r}$ . The  $2 + q_r$  involved fields have then:

$$K_r = -n_r \ln \left( \left( M_{1_r} + M_{1_r}^{\dagger} \right) \left( M_{2_r} + M_{2_r}^{\dagger} \right) - \sum_{a_r} \left( Z_{a_r} + Z_{a_r}^{\dagger} \right)^2 \right)$$

Derendinger, Kounnas, Petropoulos, Zwirner

The corresponding scalar manifold is in this case:

$$\mathcal{M}_r = rac{SO(2,2+q_r)}{SO(2) imes SO(2+q_r)} \supset \left(rac{SU(1,1)}{U(1)}
ight)^2$$

This behaves as 2 copies of the minimal geometry for the flatness and stability constraints, with an overall curvature given by:

$$R_r = rac{2}{n_r}$$

### Flatness and stability constraints

The structure of the flatness and stability constraints is blind to symmetry enhancements and is controlled by the minimal geometry:

$$\mathcal{M}_{\min} = rac{SU(1,1)}{U(1)} imes rac{SU(1,1)}{U(1)} imes \ \cdots$$

The crucial parameters in the constraints are the numerical coefficients  $n_i$ , related to the curvature of the basic submanifolds by  $R_i = 2/n_i$ . The curvature bound  $\sum_k R_k^{-1} > 3/2$  implies a simple restriction on the coefficients  $n_i$ :

$$\sum_k n_k > 3$$

The Goldstino cone is also entirely specified in terms of the  $n_i$ , and restricts the direction of supersymmetry breaking:

 $\Theta_i: \left\{egin{array}{l} ext{upper bound smaller than 1 if} & n_i < 3 \ & ext{lower bound larger than 0 if} & \sum_{k 
eq i} n_k < 3 \end{array}
ight.$ 

### Dilaton and volume moduli

The most relevant moduli are the dilaton S, controlling the coupling, and the global volume modulus T, controlling the size of the internal manifold. These universally occur in all models, with:

$$n_S=1\ ,\ n_T=3$$

Taking each field separately, the curvature bound is always violated. To fulfill it, one would need corrections. These should be large for S, but could be small for T.

Keeping both fields, the curvature bound is instead fulfilled. However T must dominate over S, and the Goldstino angle  $\theta$  is constrained to the range  $[0, \pi/4]$ .

This demonstrates in an extremely robust way that the scenario where S dominates over T is impossible to realize, at least in the controllable limit where both are large.

# CONCLUSIONS AND OUTLOOK

- In a generic SUGRA model with chiral and vector multiplets, there exist necessary conditions for flatness and stability that strongly constrain the geometry and the SUSY breaking direction.
- When *F* breaking dominates, the constraints are simple and rather strong. What matter is the Kähler curvature.
- The effect of an additional *D* breaking to alleviate the constraints.
   What matter is then a smaller effective Kähler curvature.